

Axions in astrophysics and cosmology

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November 2, 2020

Axion Lagrangian

Axions are pseudoscalars which interact with SM particles as

$$\begin{aligned}\mathcal{L}_{\text{ax}} = & \frac{1}{2} \partial_\mu a \partial^\mu a - \xi \frac{a}{f_a} \frac{g^2}{32\pi^2} \tilde{G}_{\mu\nu}^a G^{\mu\nu a} + \\ & + \frac{g_a}{2m} \bar{\Psi} \gamma^\mu \gamma^5 \Psi \partial_\mu a - \frac{g_{a\gamma}}{4} a \tilde{F}^{\mu\nu} F_{\mu\nu} - \frac{1}{2} m_a^2 a^2\end{aligned}$$

where f_a is a typical energy scale, the Peccei-Quinn scale, and m_a is the axion mass

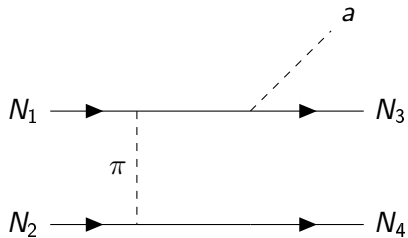
Axions could solve

- ▶ The strong CP problem
- ▶ The origin of dark matter
- ▶ Some stellar anomalies

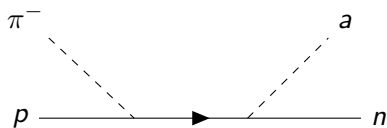
Axion production channels

Axions coupled to nucleons

NN bremsstrahlung



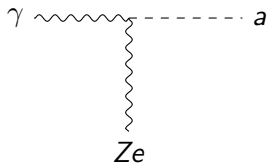
Pion-axion conversion



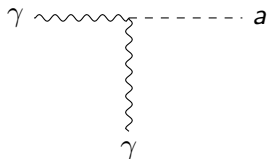
Axion production channels

Axions coupled to photons

Primakoff conversion



Inverse Decay



The energy-loss argument

G. Raffelt, Lect. Notes Phys. **741** (2008)

Stars produce axions which escape, draining energy from the core



Axions affect strongly the SN explosion if

$$\epsilon_a > 10^{19} \text{ erg g}^{-1} \text{ s}^{-1}$$

and HB star evolution if

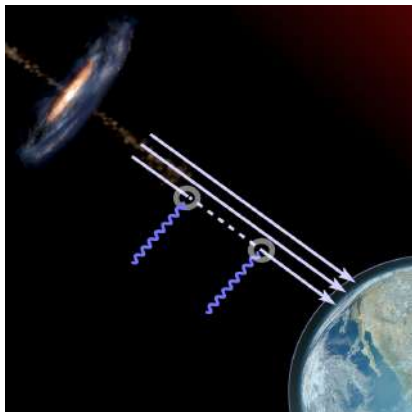
$$\epsilon_a > 10 \text{ erg g}^{-1} \text{ s}^{-1}$$

ALP-photon oscillations

The ALP-photon interaction

$$\mathcal{L} = g_{a\gamma} a \mathbf{E} \cdot \mathbf{B}$$

allows for photon-ALP conversions in the Galactic or stellar B fields

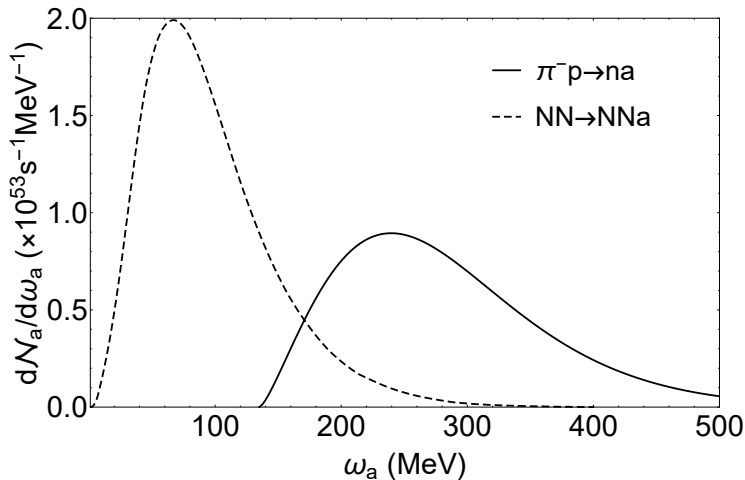


Enhanced Supernova Axion Emission and its Implications

In collaboration with B. Fore, M. Giannotti, A. Mirizzi and S. Reddy
arXiv:2010.02943 [hep-ph]

Pions in SN1987A

Axions are produced by pion-axion conversion ($Y_\pi \sim 1\%$)



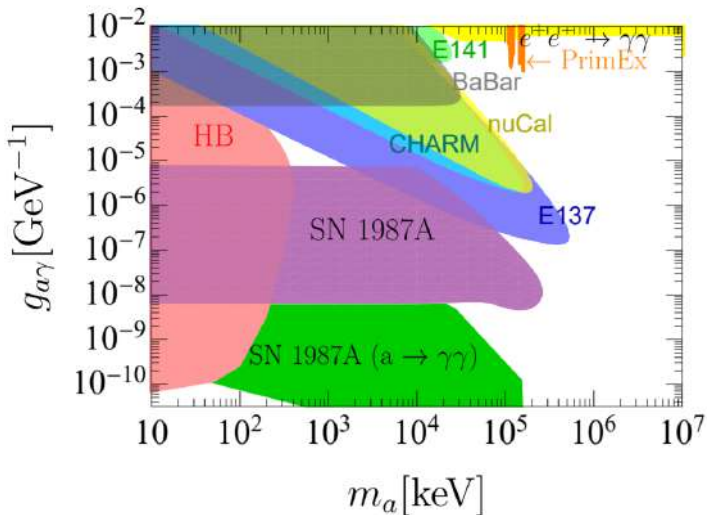
The SN cooling bound is $\times 2$ stronger

Heavy axion-like particles and core-collapse supernovae: constraints and impact on the explosion mechanism

In collaboration with G. Lucente, T. Fischer, M. Giannotti and A. Mirizzi
arXiv:2008.04918 [hep-ph]

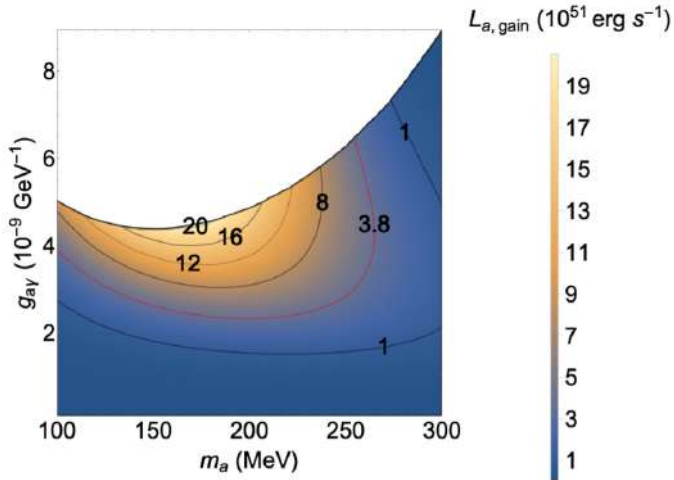
Axion-Like Particles from SNe: SN1987A bound

ALP are produced via Primakoff conversion and Inverse Decay



Can ALP revitalize the SN shock?

Massive ALP could decay inside the SN revitalizing the shock



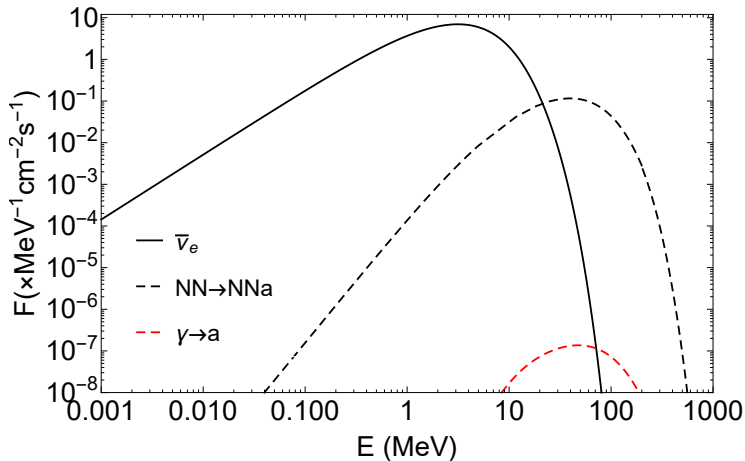
Energy deposited at $t_{\text{pb}} = 0.3 \text{ s}$, the red line indicates where the ALP deposit the same energy as neutrinos

Bounds on axion-like particles from the diffuse supernova flux

In collaboration with F. Calore, M. Giannotti, J. Jaeckel and A. Mirizzi
arXiv:2008.11741 [hep-ph]

DSNALPB

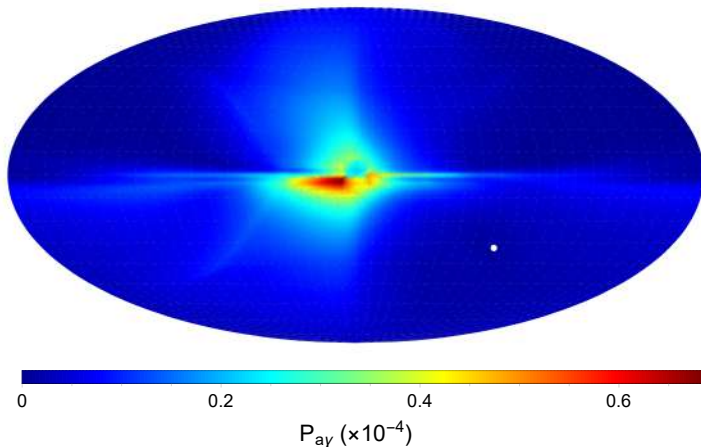
In analogy to neutrinos, the DSNALPB is created by all past SNe



DSNALPB with $g_{ap} = 1.2 \times 10^{-9}$ and $g_{a\gamma} = 5.3 \times 10^{-12} \text{ GeV}^{-1}$

ALP conversion into photons

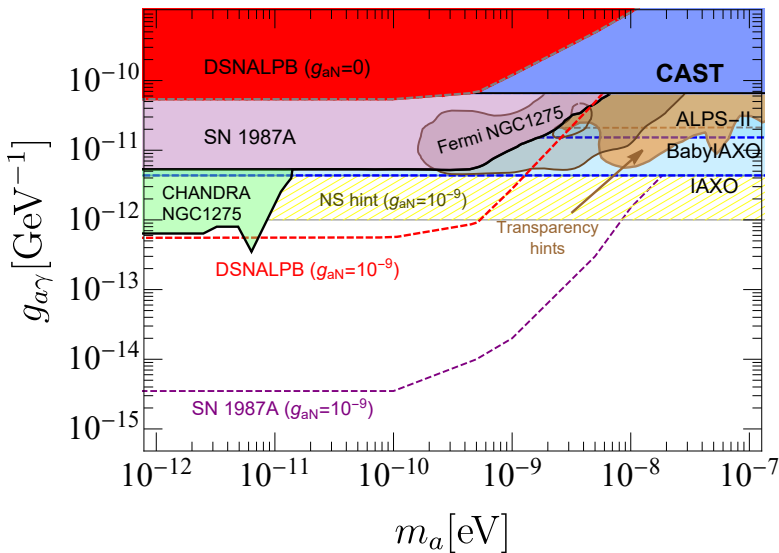
The Galactic magnetic field will convert into photons both the DSNALPB and the point-like ALP flux from SN1987A (white dot)



Conversion probability for $m_a \ll E = 50 \text{ MeV}$, $g_{a\gamma} = 3 \times 10^{-13} \text{ GeV}^{-1}$

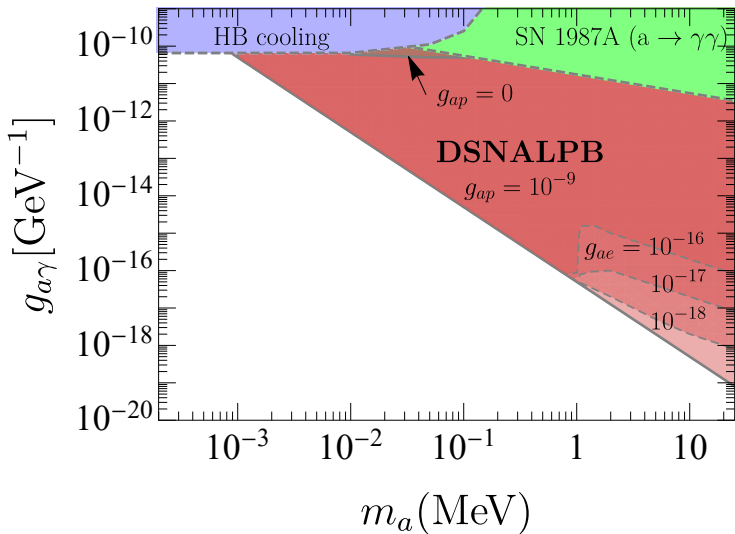
Bound on neV-scale ALPs

Diffuse -ray background constrained by Fermi-LAT



Bound on MeV-scale ALPs

Diffuse -ray background constrained by Fermi-LAT+COMPTEL



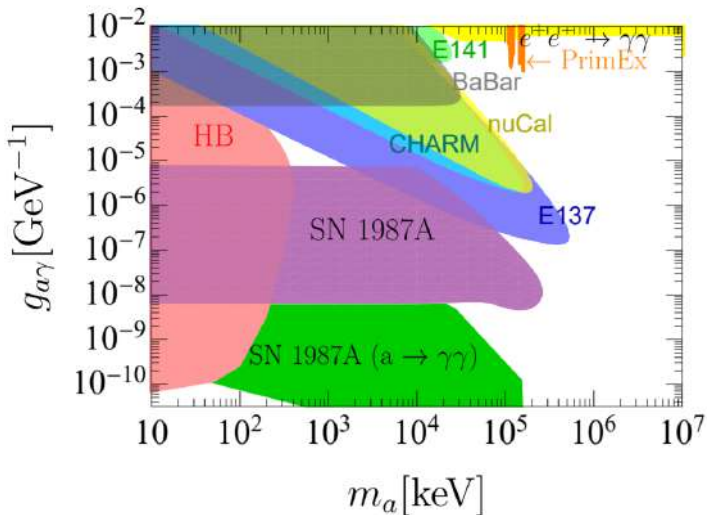
Constraints on the coupling with photons of heavy axion-like-particles from Globular Clusters

In collaboration with O. Straniero, B. Döbrich, M. Giannotti, G. Lucente and A. Mirizzi

arXiv:2004.08399 [hep-ph]

HB star bound on heavy ALPs

A small region is unconstrained: the “cosmological triangle”

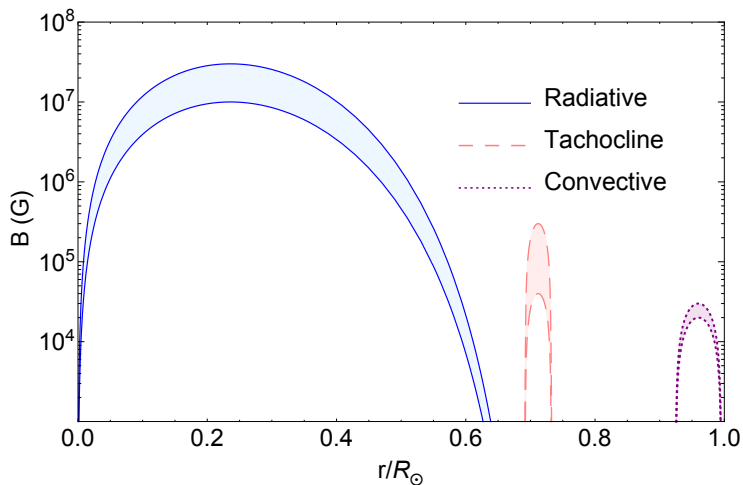


Production of axion-like particles from photon conversions in large-scale solar magnetic fields

In collaboration with E. Guarini, J. Galan, M. Giannotti and A. Mirizzi
arXiv:2010.06601 [hep-ph]

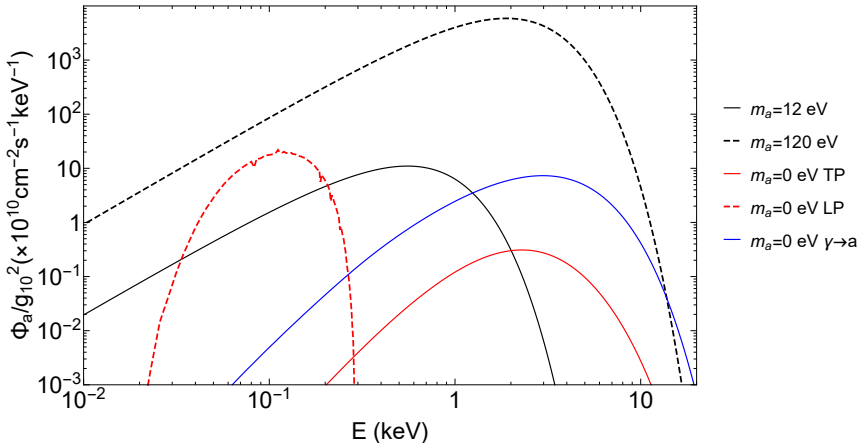
A new production channel

Photons convert into axions in the solar magnetic field



Comparison of the fluxes

New phenomenology is related to this flux



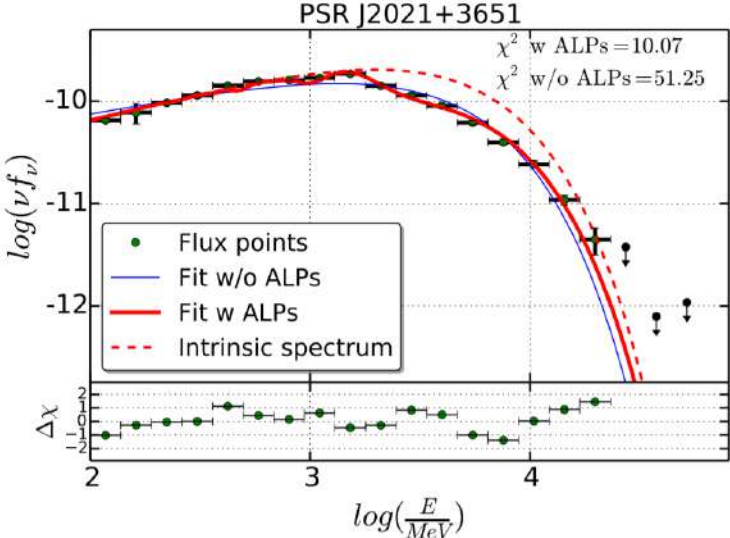
Reconciling hints on axion-like-particles from high-energy gamma rays with stellar bounds

In collaboration with G. A. Pallathadka, F. Calore, M. Giannotti,
D. Horns, J. Majumdar, A. Mirizzi, A. Ringwald, A. Sokolov and F. Stief
arXiv:2008.08100 [hep-ph]

Updated PSR hint on ultralight ALPs

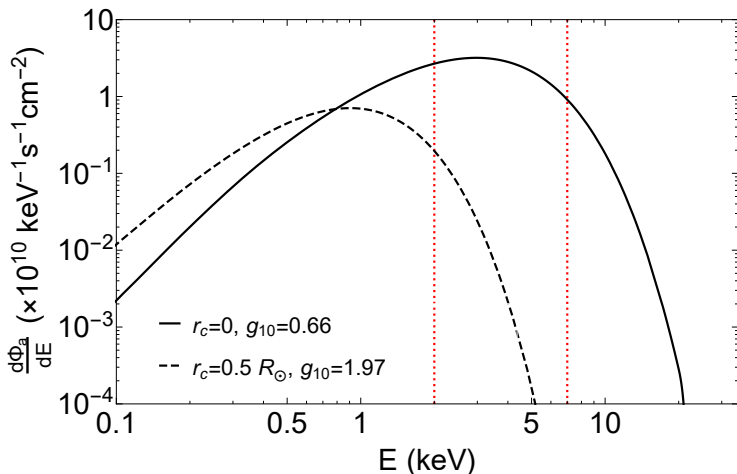
J. Majumdar, F. Calore and D. Horns, JCAP **04** (2018), 048

ALP with $m_a = 4 \text{ neV}$ and $g_{a\gamma} = 1.97 \times 10^{-10} \text{ GeV}^{-1}$



ALPs at CAST

The CAST bound evaded by environmental effects: $g_{a\gamma}(\eta)$

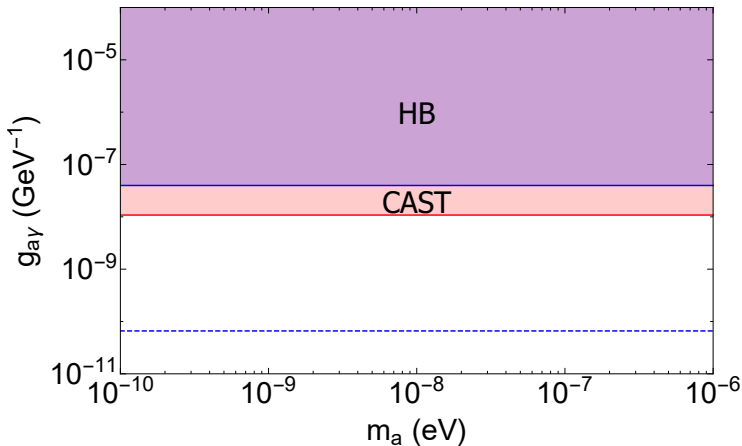


The new bound is $g_{a\gamma} < 2.37 \times 10^{-9} \text{ GeV}^{-1}$.

HB and SN bound

If $r_c = 0.5 R_\odot$, assuming that $g_{a\gamma}(\rho)$, then $\rho_c = 1.3 \text{ g cm}^{-3}$

The HB bound is relaxed to $g_{a\gamma} < 4 \times 10^{-8} \text{ GeV}^{-1}$

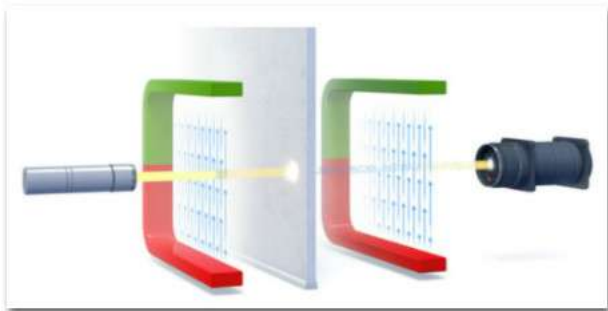


The SN bound disappears because $\rho \sim 10^{14} \text{ g cm}^{-3}$

ALPS II

R. Bähre *et al.*, JINST **8** (2013), T09001

The future LSTW experiment, ALPS II, can easily detect ALPs with $m_a \ll 10^{-4}$ eV and $g_{a\gamma} = 1.97 \times 10^{-10} \text{ GeV}^{-1}$



Schools attended:

- ▶ Winter School on Multi-Messenger Astrophysics, Asiago (Italy), 14 - 23 Jan. 2020

Conferences:

- ▶ Workshop “Axion cosmology,” Munich (Germany), 17-28 Feb. 2020
- ▶ New Frontiers in Theoretical Physics - Convegno nazionale di fisica teorica, e-conference, 27-29 May 2020
- ▶ DESY Virtual Theory Forum, e-conference, 22-25 Sep. 2020
- ▶ IBS-ICTP Workshop on Axion-Like Particles, e-conference, 21-23 Oct. 2020

Scientific collaborations:

- ▶ Scientific collaboration with O. Straniero
Osservatorio Astronomico dell’Abruzzo, Teramo (Italy), Dec. 2019

Exams: Completed

Talks:

- ▶ *“Improved axion emissivity from a supernova and the SN1987A bound”*
Department of Physics, Bari (Italy), 17 Dec. 2019
- ▶ *“Stellar bounds on axions and ALPs”*
Munich Institute for Astro- and Particle Physics (MIAPP),
Munich (Germany), 26 Feb. 2020
- ▶ *“Constraints on the coupling with photons of heavy axion-like-particles from Globular Clusters”*
online talk for the IAXO, MADMAX and ALPS collaborations,
28 May 2020
- ▶ *“Reconciling hints on axion-like-particles from high-energy gamma rays with stellar bounds”*
online talk for the Virtual Axion Institute, hosted by Kai
Schmitz and Valerie Domcke, 1 Sep. 2020
- ▶ *“Bounds on axion-like particles from the diffuse supernova flux”*
online talk for the DESY Virtual Theory Forum 2020, 22 Sep.
2020

Thanks for your attention