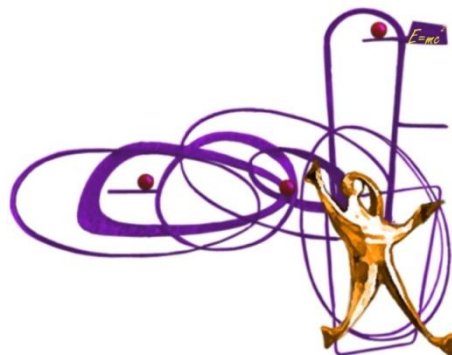


# Studio delle sezioni d'urto di reazioni indotte dai neutroni nell'ambito Astrofisico presso la facility n\_TOF al CERN

**Dottorato di ricerca in Fisica XXXI ciclo**

**Università degli studi di Bari Aldo Moro, Bari, 9 Novembre 2015**



Dottoranda:

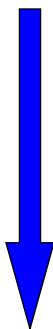
**Lucia Anna Damone**

- Motivazioni scientifiche
- La facility n\_TOF
- Il lavoro di Tesi
- Programma di lavoro per i tre anni
- Conclusioni

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>100 Ricercatori

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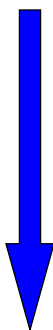
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$(n,f)$ ,  $(n,cp)$ ,  $(n,\gamma)$

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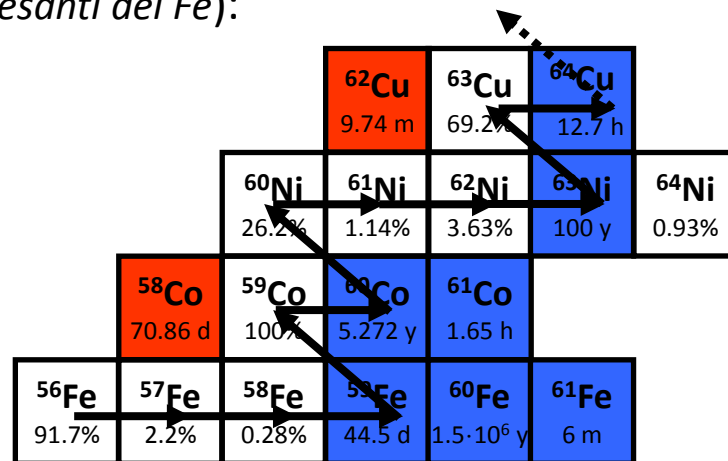
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## Astrofisica Nucleare (*nucleosintesi stellare elementi più pesanti del Fe*):

- processi di **cattura neutronica** lenti (AGB stars) e veloci (Supernovae) e successivi **decadimenti  $\beta$** ;
- Produzione degli elementi più leggeri nella BBN

**Sezioni d'urto di cattura** fondamentali per capire l'abbondanza degli elementi pesanti nell'Universo e per modelli sull'**evoluzione** delle **stelle** e delle **galassie**.

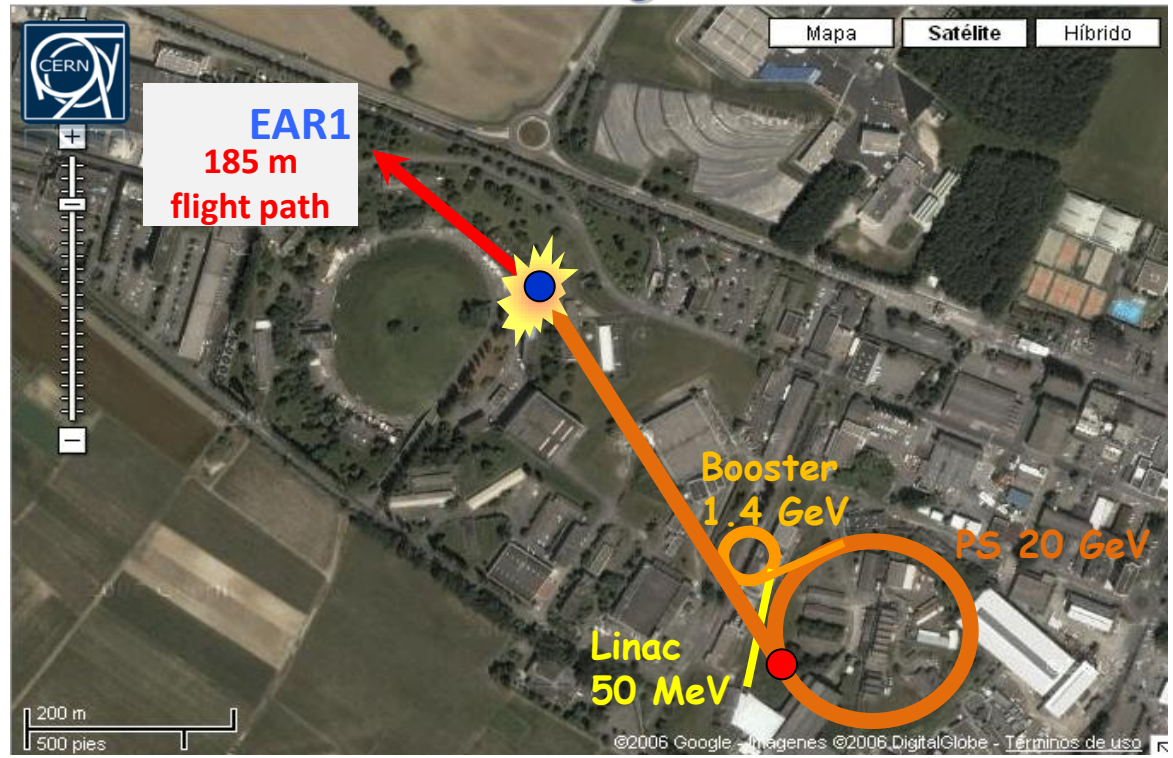


## Applicazioni (energia da **fissione**, **fusione**, **Medicina Nucleare**, **Neutron imaging**)

Sezioni d'urto di **cattura** e **fissione** importanti per :

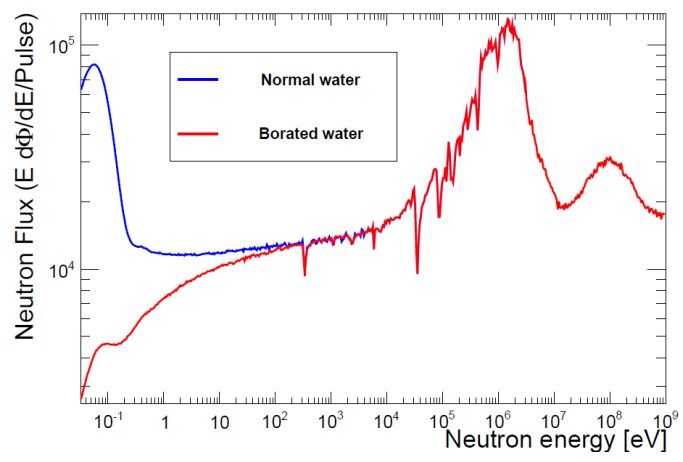
- migliorare sicurezza ed efficienza reattori nucleari attuali (termici) o sviluppare nuovi sistemi per **produzione di energia e trasmutazione scorie** (Gen IV e ADS);
- Stima dei danni da radiazione nei futuri reattori a fusione
- **Medicina nucleare**: nuove terapie, produzione di radioisotopi

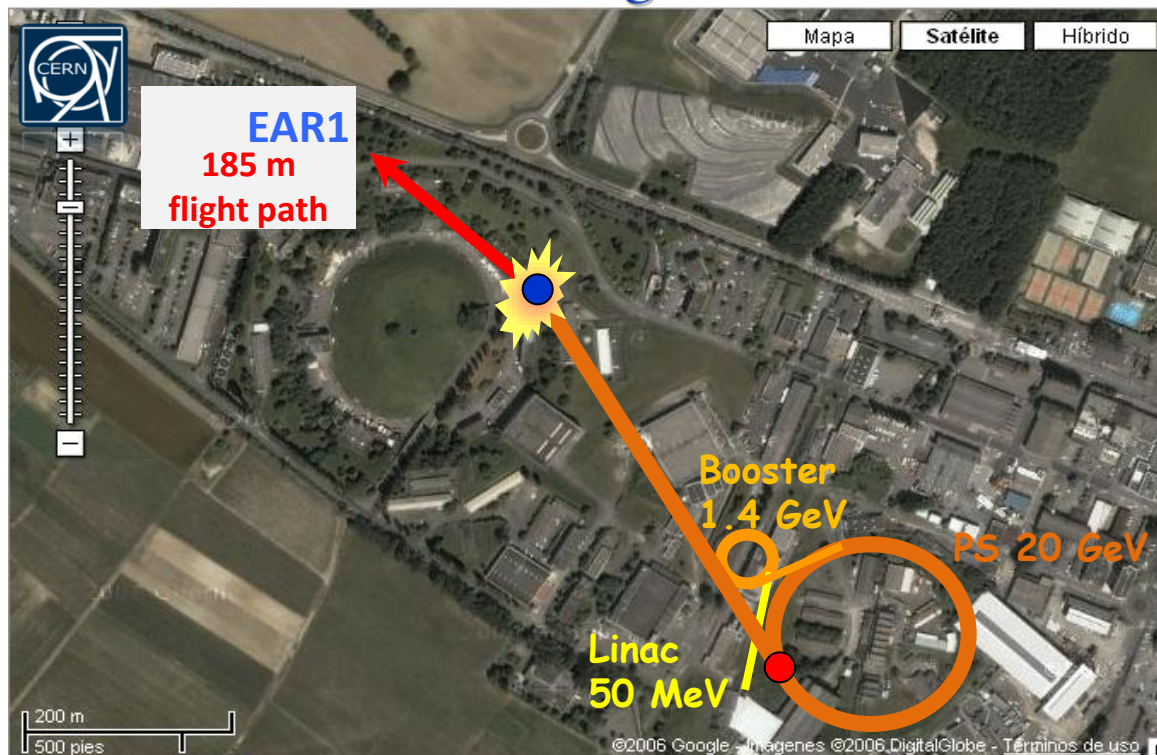
# La facility n\_TOF



$p(20 \text{ GeV}/c) \rightarrow Pb \rightarrow n$   
 (~360 neutrons per proton)

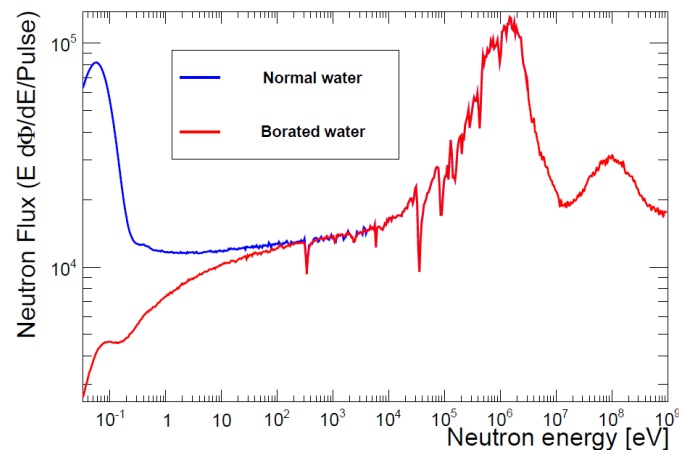
**Fino al 2014** sala sperimentale  
**EAR1** a 185 m dal target di  
 spallazione.





$p(20 \text{ GeV}/c) \rightarrow \text{Pb} \rightarrow n$   
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**EAR1** a **185 m** dal target di  
 spallazione.



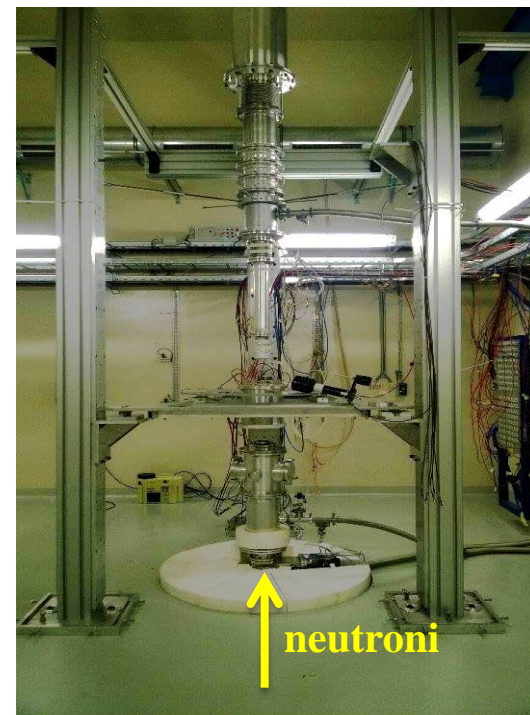
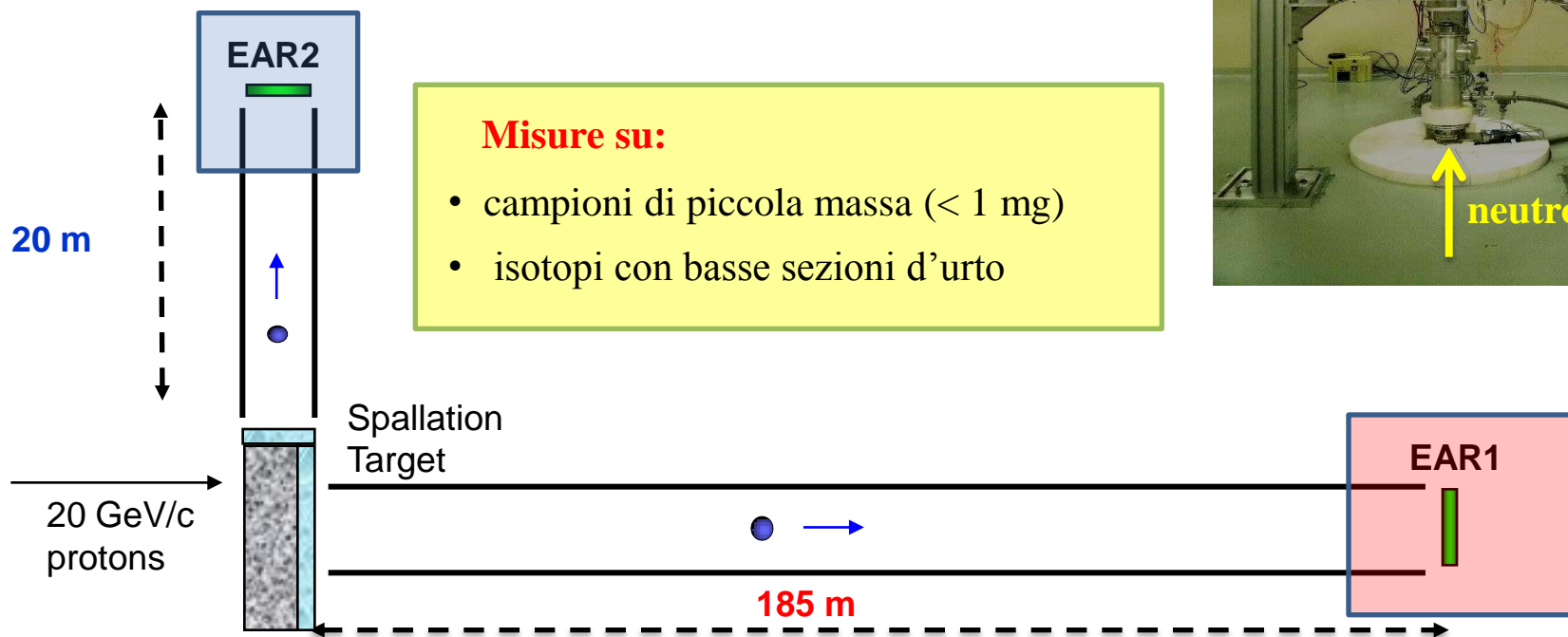
## Principali caratteristiche:

- **Flusso istantaneo di neutroni estremamente alto**  
 ( $10^5 \text{ n}/\text{cm}^2/\text{pulse}$ ).
- **Ampio spettro energetico** (dal termico fino al GeV)
- **Alta risoluzione** ( $\Delta E/E=10^{-4}$ )

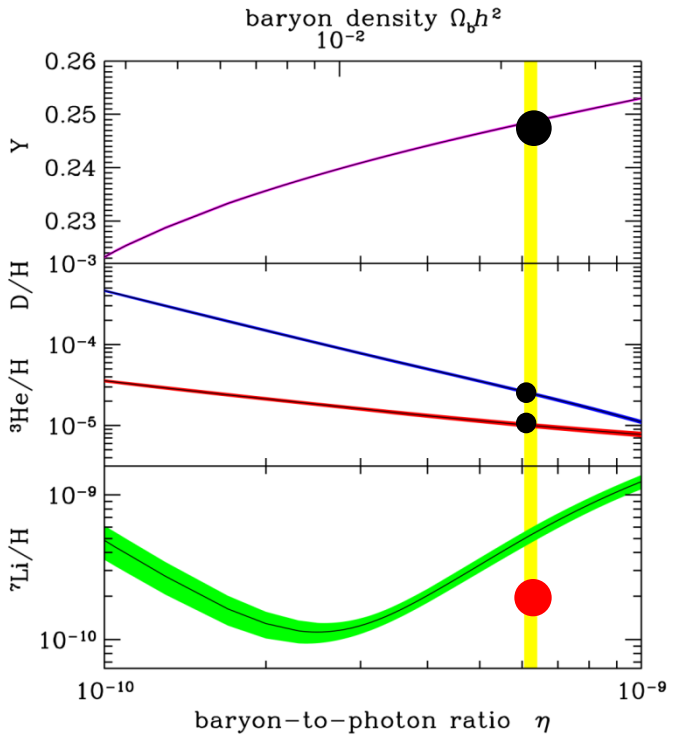


Dal 2014 seconda area sperimentale (EAR2) a **20 m** dal target di spallazione.

- **Alto flusso** ~ fattore 30 più alto rispetto ad EAR1.
- **Flight path più breve** → tempo di volo 10 volte più piccolo.
- Aumento di un **fattore 300** nel rapporto **signal/background** per gli isotopi radioattivi!



# Il problema del Litio Cosmologico



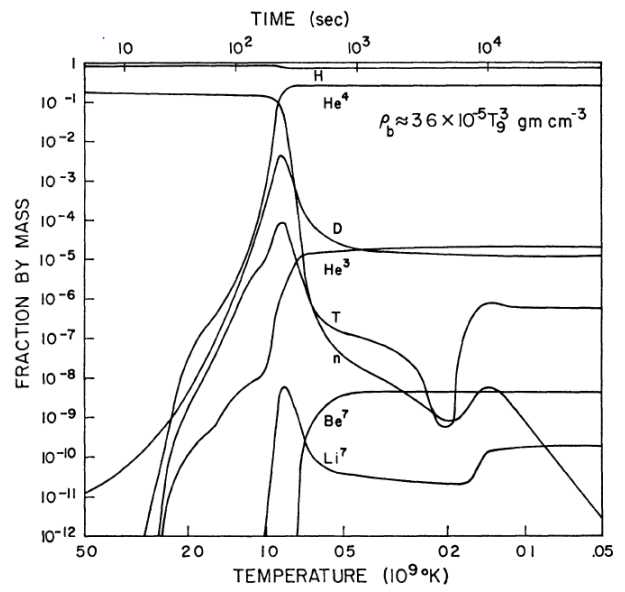
**<sup>7</sup>Li sovrastimato di un fattore 3-5**

**Possibile spiegazione:** teoria standard → 95% del <sup>7</sup>Li prodotto dal decadimento del <sup>7</sup>Be (tempo di dimezzamento 53.2 giorni).



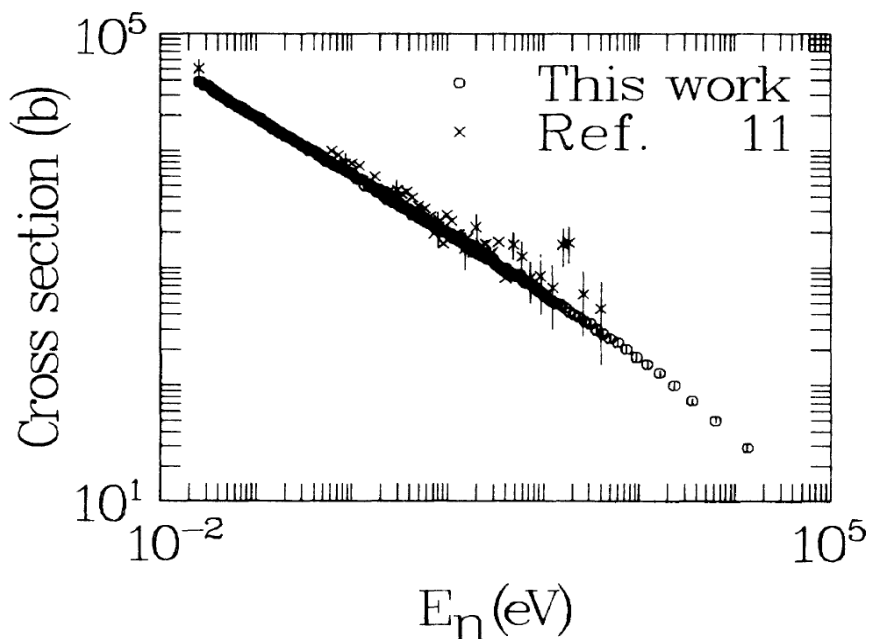
**Abbondanza del <sup>7</sup>Li legata alla produzione e distruzione del <sup>7</sup>Be.**

Reazioni indotte da **protoni**, **deutoni** e **<sup>3</sup>He non responsabili** della distruzione del <sup>7</sup>Be.





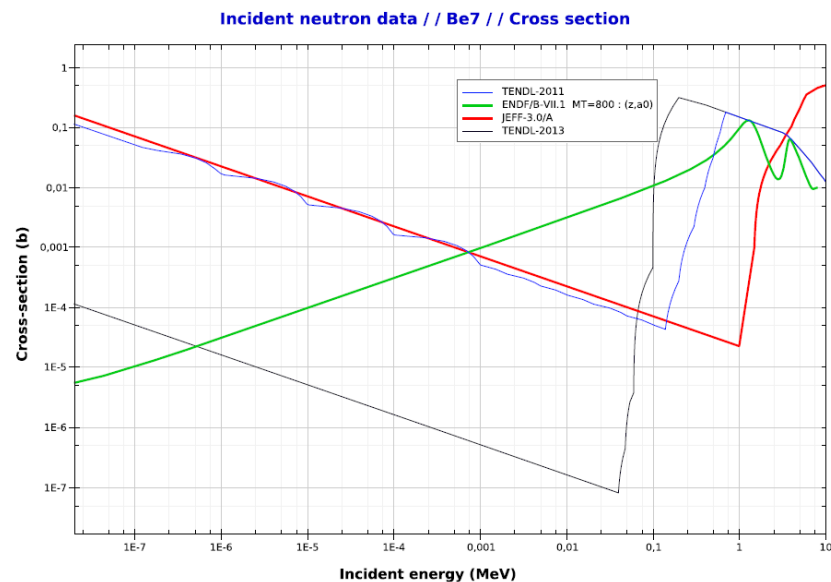
→ dal **termico** a **13.5 keV**  
(P. Koehler et al, Los Alamos 1988)



**Dottorato di ricerca:** studio della reazione  ${}^7\text{Be} (n,p) {}^7\text{Li}$  nel range di interesse del problema (**20 - 100 keV**)



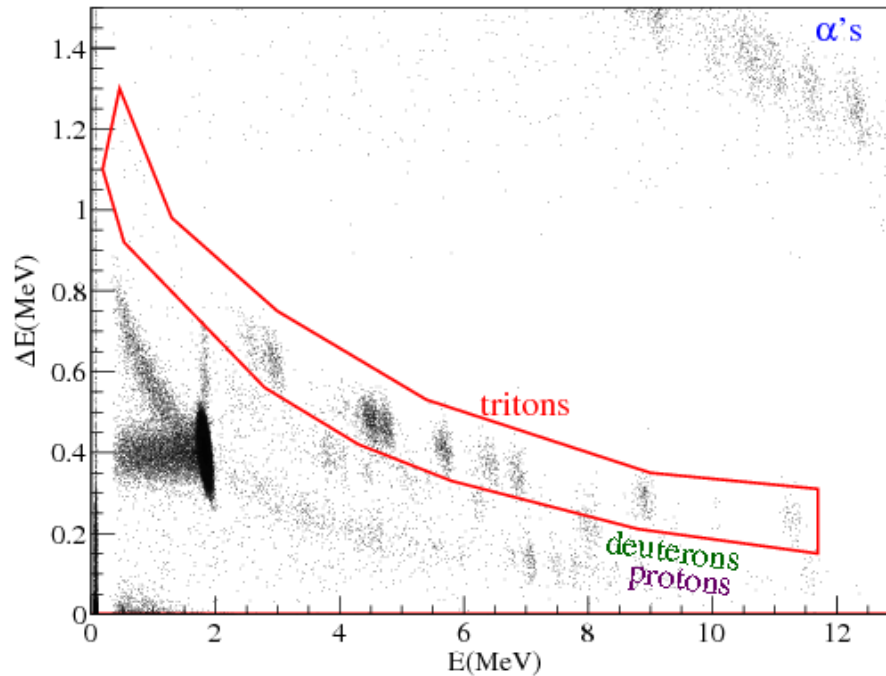
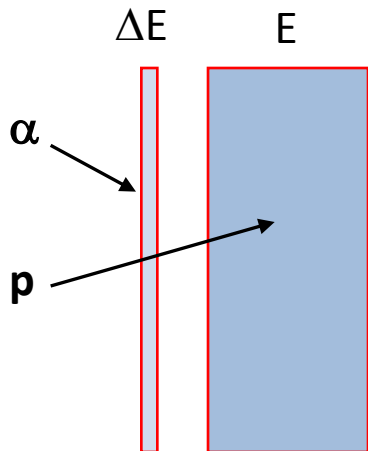
→ **valutazioni teoriche**  
(P. Bassi et al., ISPRA '60)



Nell'ambito del **lavoro di tesi** mi sono occupata del test dell'apparato usato nella  ${}^7\text{Be} (n,\alpha) {}^4\text{He}$

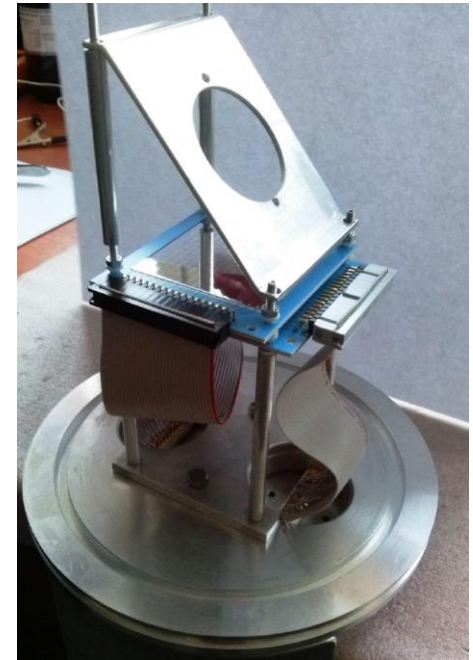
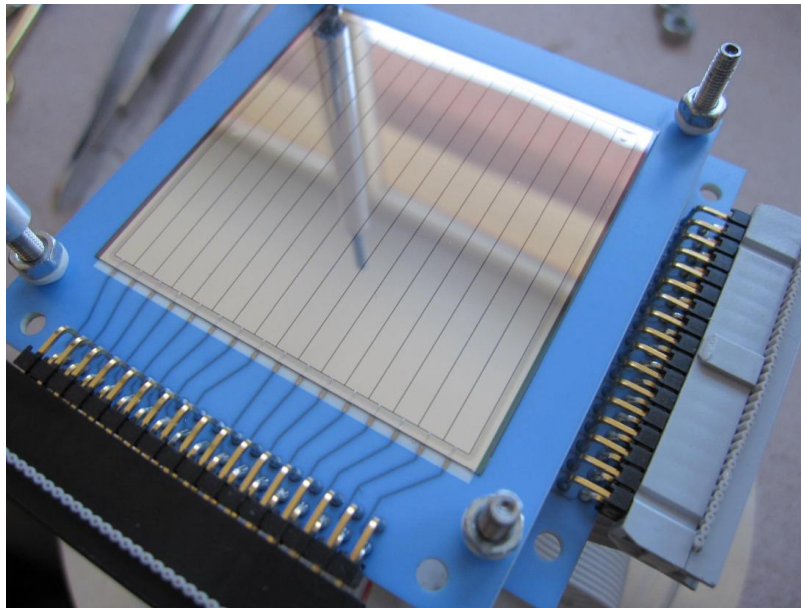
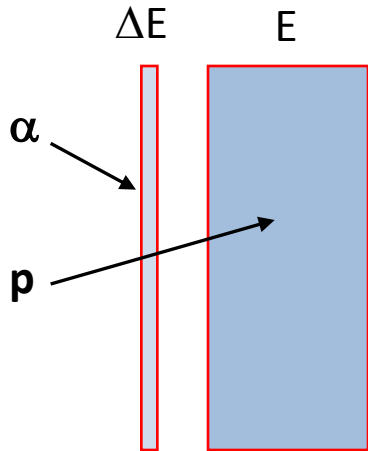
Tecnica per lo studio delle reazioni di particella carica indotte dai neutroni [ $(n,p)$   $(n,\alpha)$ ...]:  
**Silicon Telescope  $\Delta E$ -E.**

Due **rivelatori a strip di Silicio** (16 strips,  $5 \times 5 \text{ cm}^2$ , spessore **20 e 300  $\mu\text{m}$** )



Tecnica per lo studio delle reazioni di particella carica indotte dai neutroni [(n,p) (n, $\alpha$ )...]:  
Silicon Telescope  $\Delta E$ -E.

Due rivelatori a strip di Silicio (16 strips,  $5 \times 5 \text{ cm}^2$ , spessore 20 e 300  $\mu\text{m}$ )



## Primo anno

- montaggio e test dell'apparato sperimentale ;
- presa dati dell'esperimento;
- simulazioni Monte Carlo per la stima dell'efficienza dell'apparato;

## Secondo anno

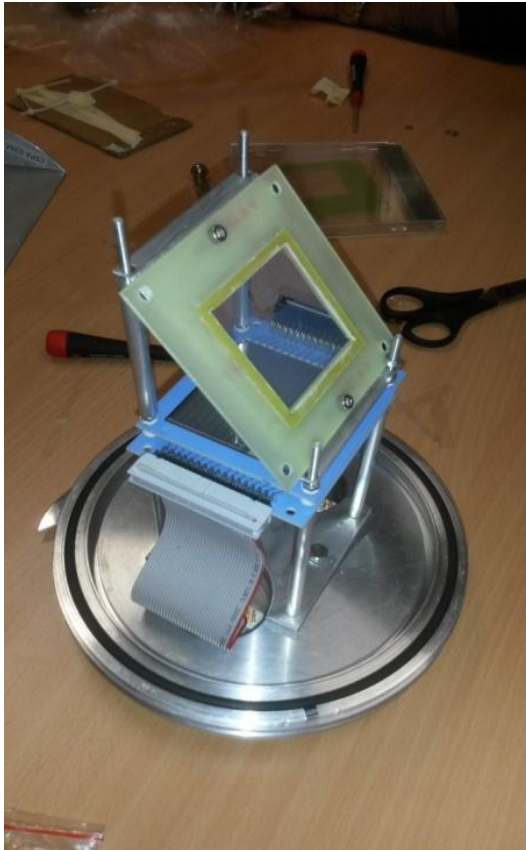
- Analisi dei dati per l'estrazione della sezioni d'urto;
- Studio di modelli di BBN e dell'impatto delle nuove misure ad n\_TOF sui calcoli dell'abbondanza primordiale;

## Terzo anno

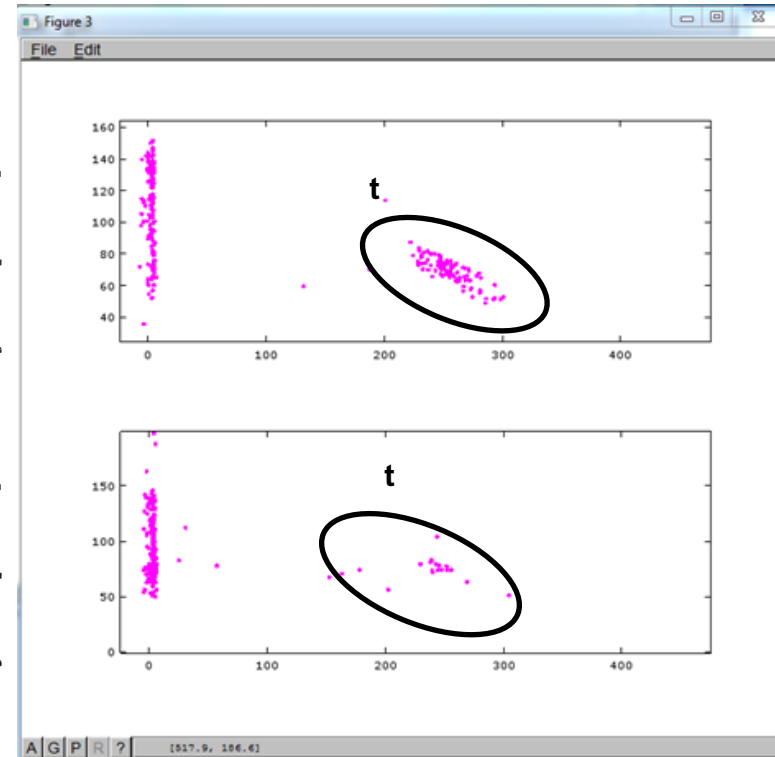
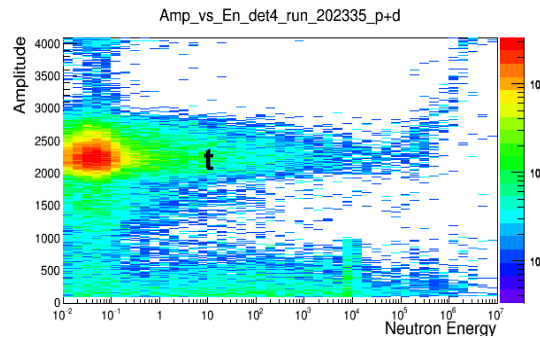
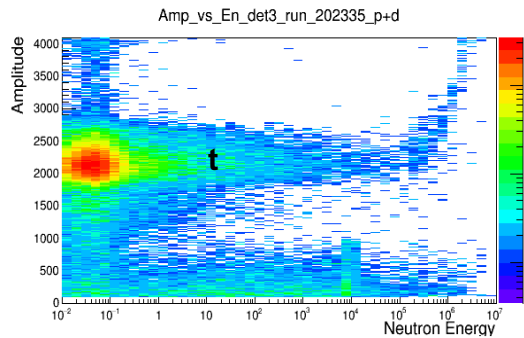
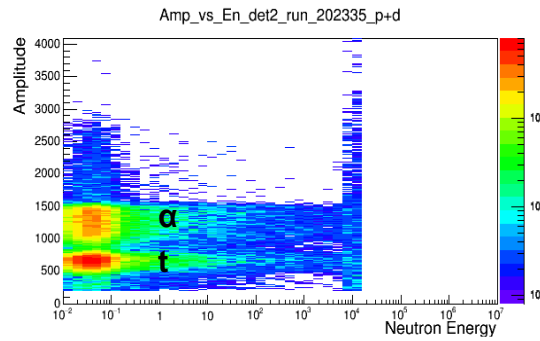
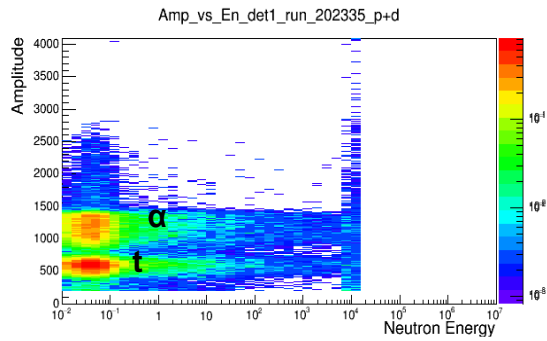
- Studio di una reazione di cattura neutronica di interesse astrofisico (Ca, Y, etc...)

**Grazie per l'attenzione**

- Convertitore di **LiF** spesso  **$105 \mu\text{g}/\text{cm}^2$**  (LNS) ( ${}^6\text{Li}(n,t){}^4\text{He}$ )
- **Particelle  $\alpha$**  da 2.05 MeV sono fermate nel rivelatore  $\Delta E$
- **Tritoni** da 2.73 MeV perdono  $\sim 1$  MeV nel  $\Delta E$







**La regione dei tritoni è chiaramente visibile**