

Second year Ph.D. research activity

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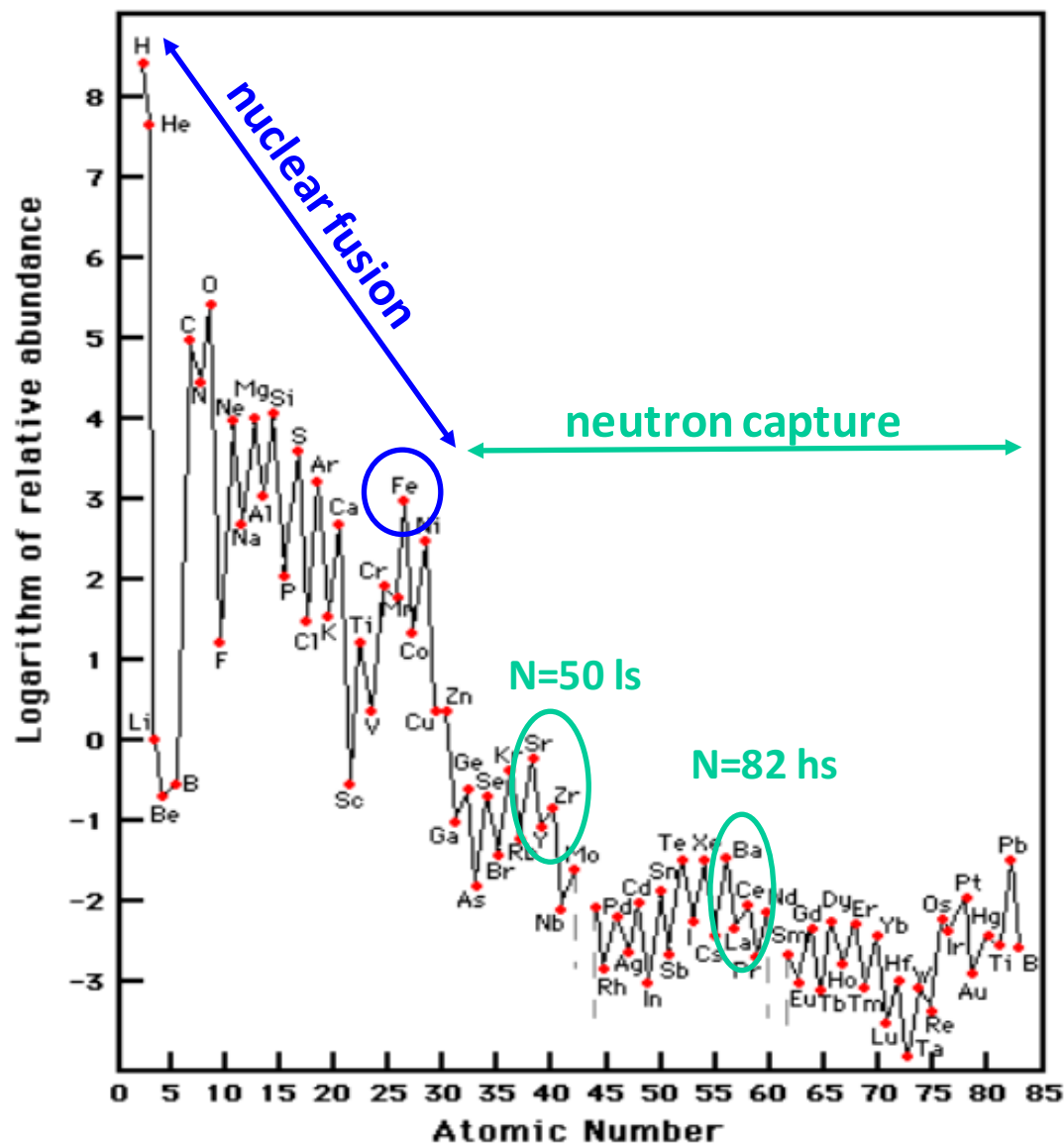
Main topics

- ❖ I stayed at CERN as an INFN Associate (“simil-fellow”)
- ❖ I concluded my analysis on the ${}^7\text{Be}(n,p){}^7\text{Li}$ cross section measurement for the Cosmological Lithium Problem
- ❖ I was responsible for the measurement of the neutron capture cross section of ${}^{89}\text{Y}$ and ${}^{88}\text{Sr}$ performed in EAR1 at $\overline{n_TOF}$





^{89}Y and ^{88}Sr neutron capture measurement



neutron capture processes

- slow neutron capture process (s-process)
 $\tau_n > \tau_\beta$ and $n_n \sim 10^8 \text{ cm}^{-3}$
- rapid neutron capture process (r-process)
 $\tau_n < \tau_\beta$ and $n_n \sim 10^{22} \text{ cm}^{-3}$

s-process bottlenecks

their neutron-capture cross sections are lower than those of neighboring nuclei. As a result, they act as bottlenecks on the neutron-capture path.



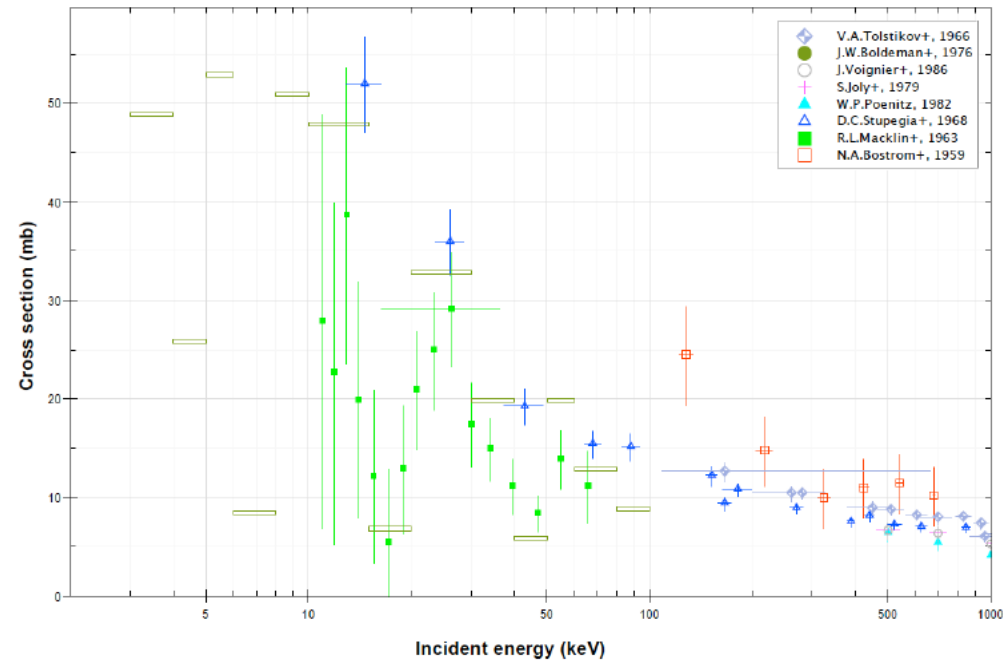
^{89}Y and ^{88}Sr neutron capture measurement

The ^{88}Sr and ^{89}Y abundances in stars easy to derive thanks to a large number of strong lines in spectra.

constrain **stellar models** and **astrophysical scenarios**.

^{89}Y hydride offers advantages as a **moderator** for high temperature thermal nuclear reactors.

It retains its relatively high content of hydrogen at very high temperatures, between $850\text{ }^{\circ}\text{C}$ to $1150\text{ }^{\circ}\text{C}$.

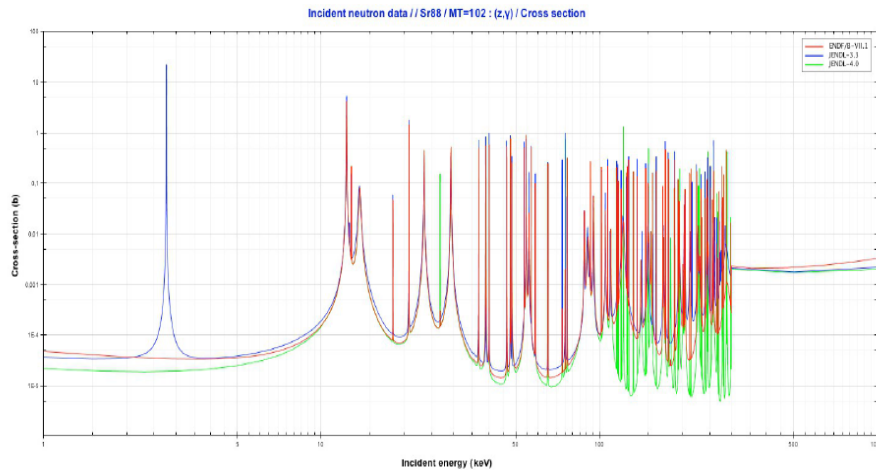




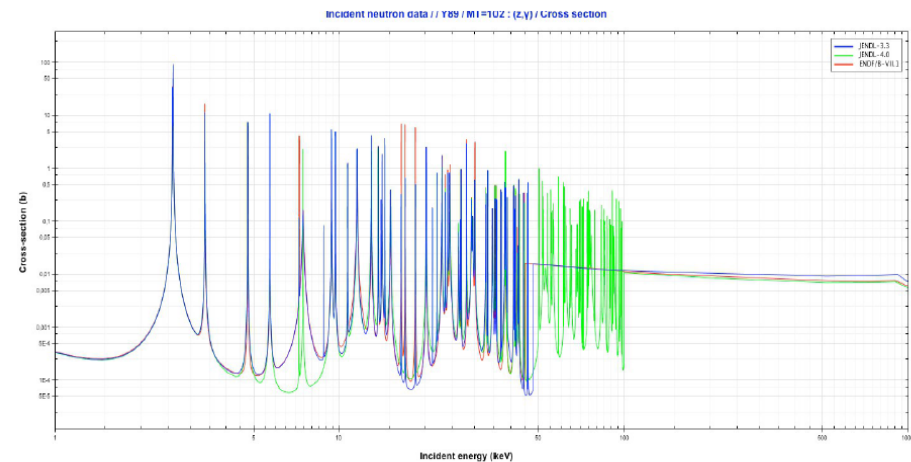
^{89}Y and ^{88}Sr neutron capture measurement

The status of the **experimental data** is also reflected in the quality of the cross sections in the evaluated data libraries.

$^{88}\text{Sr}(n,\gamma)$



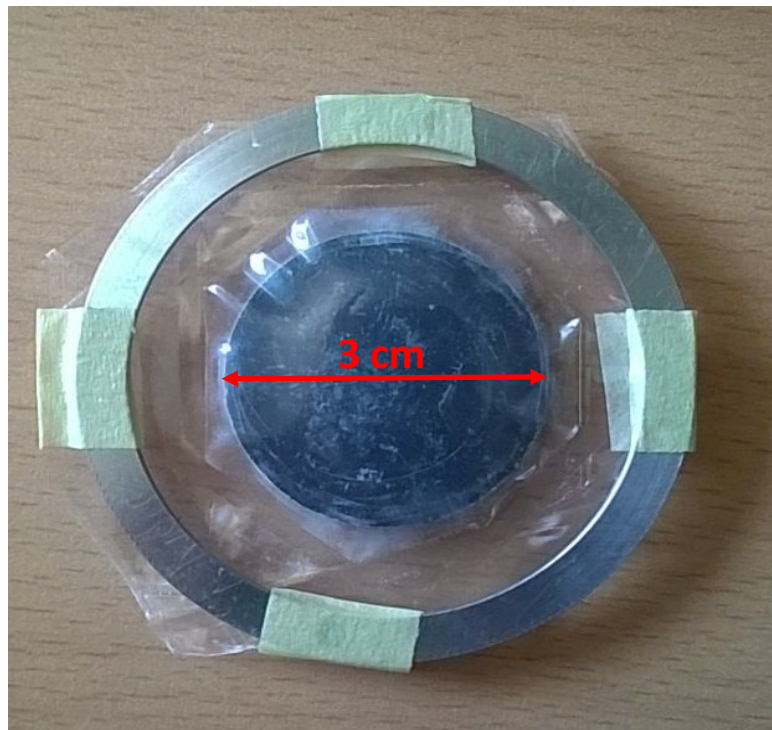
$^{89}\text{Y}(n,\gamma)$





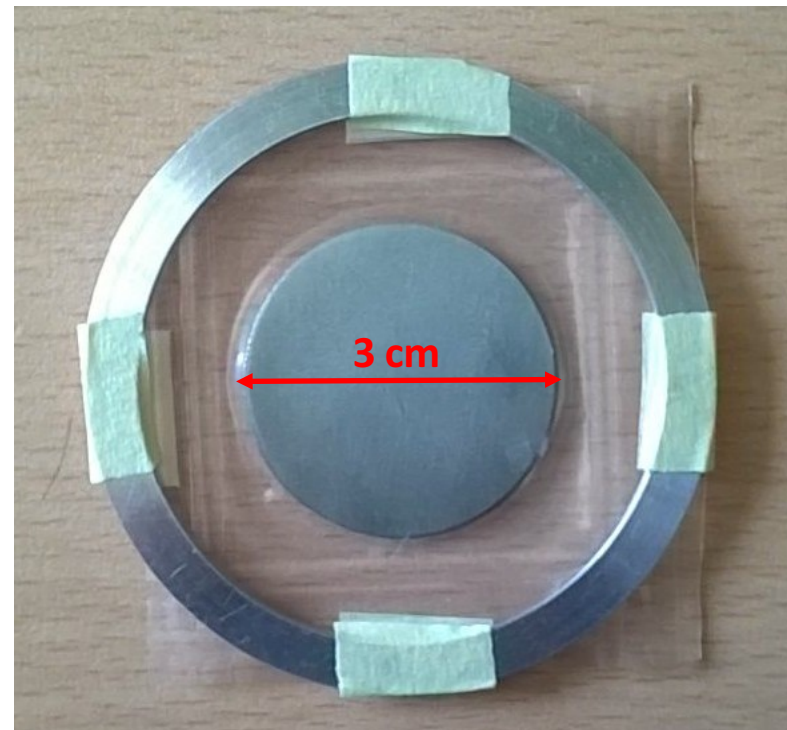
^{89}Y and ^{88}Sr samples

^{88}Sr is the most abundant Sr isotope
(82% in the Sun)



**carbonate powder with an
enrichment > 99.9%**

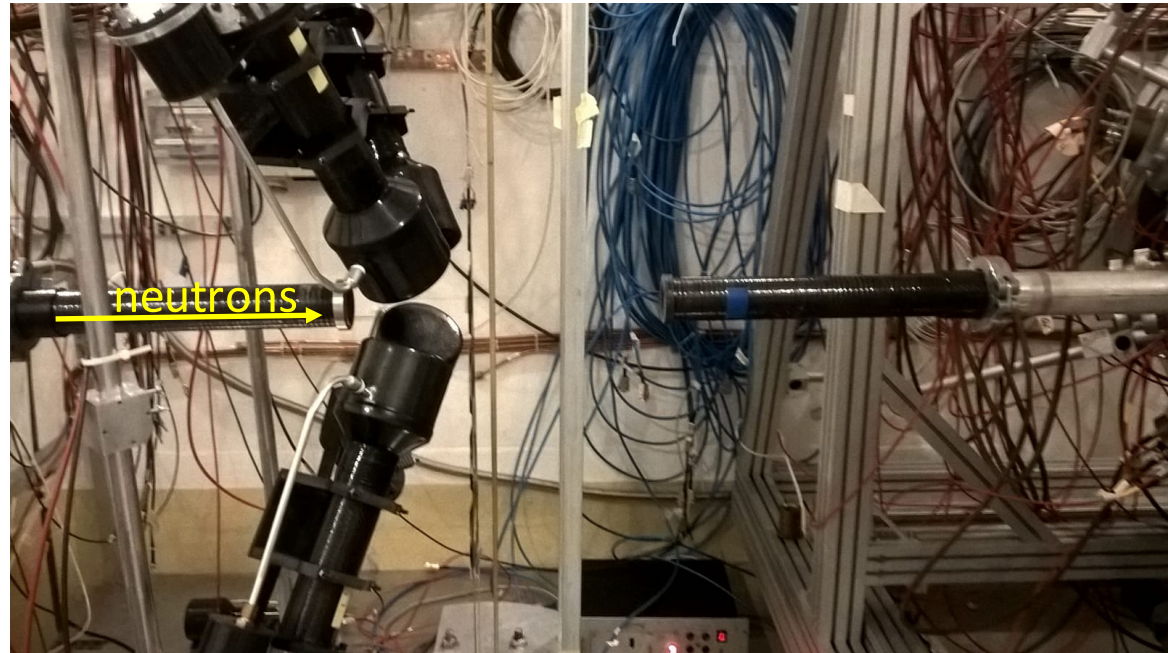
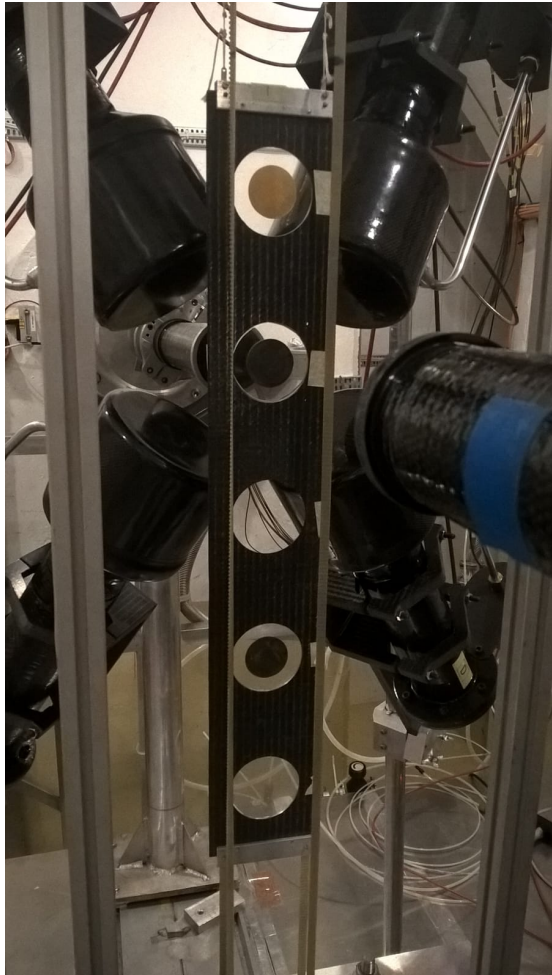
^{89}Y is the only stable isotope of Y



metal disk with a purity of 99.9%



Experimental set – up

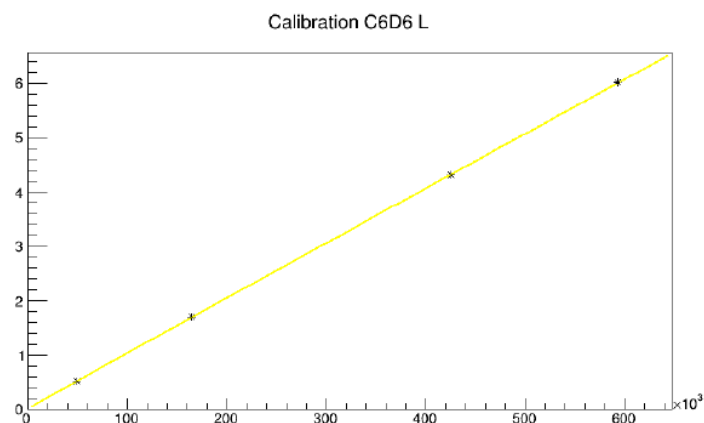
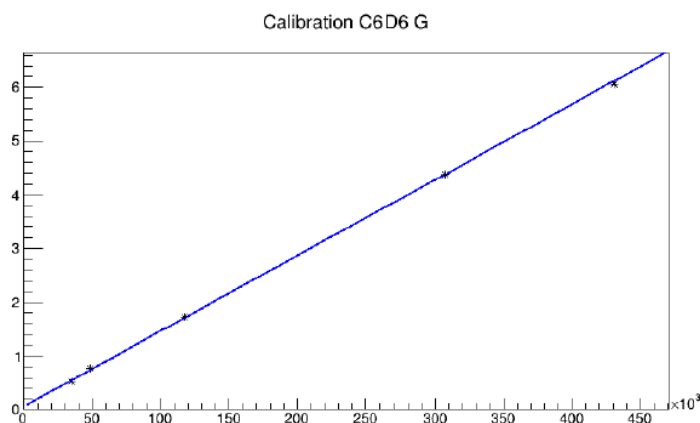
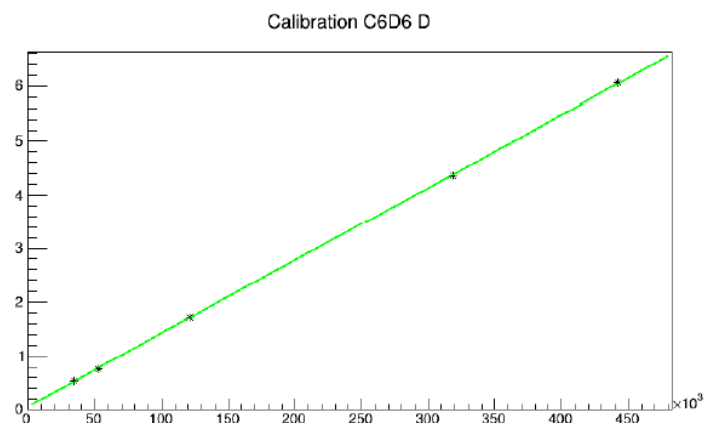
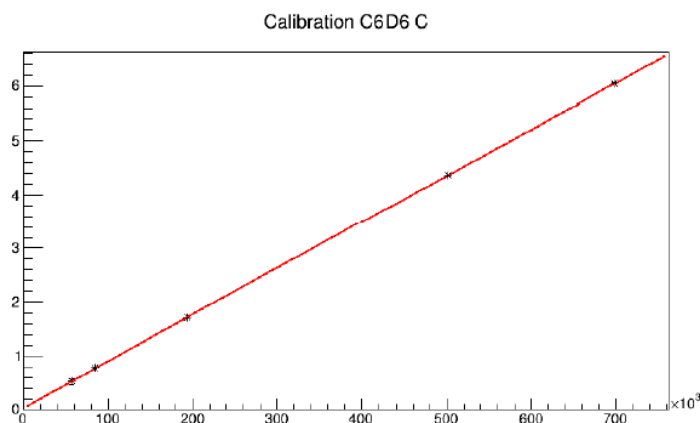


- **SiMon** a silicon based neutron beam monitor based on the ${}^6\text{Li}(n,t)\alpha$ reaction
- **4 C₆D₆ Liquid scintillators**
- **Sample Exchanger**



Energy calibration of the C6D6 detectors

Calibration γ -ray sources: ^{137}Cs , ^{88}Y , AmBe, CmC





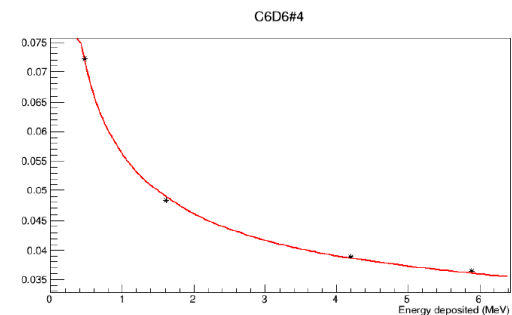
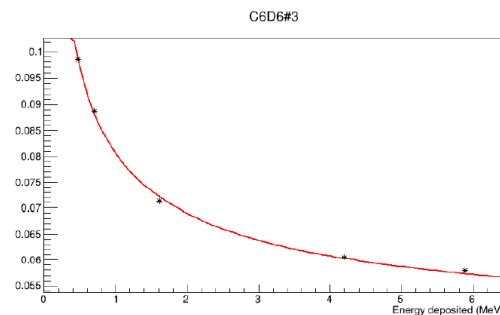
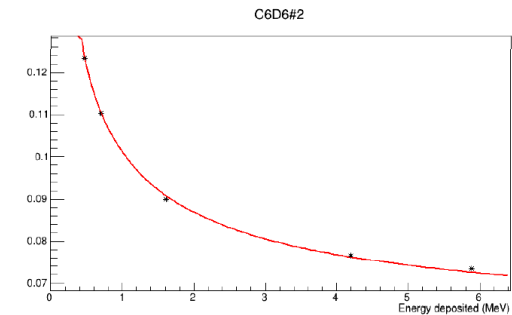
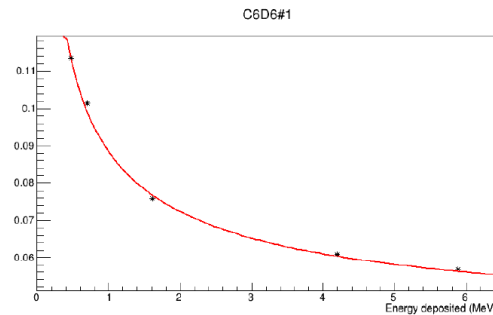
Pulse height weighting technique

The weighting functions ensure that the efficiency ϵ_γ of a detector is linear with the E_γ

The weighting function $W(E)$ is determined by minimizing the expression:

$$\sum_j \left[\int W(E') R(j; E') dE' - \alpha E_\gamma(j) \right]^2$$

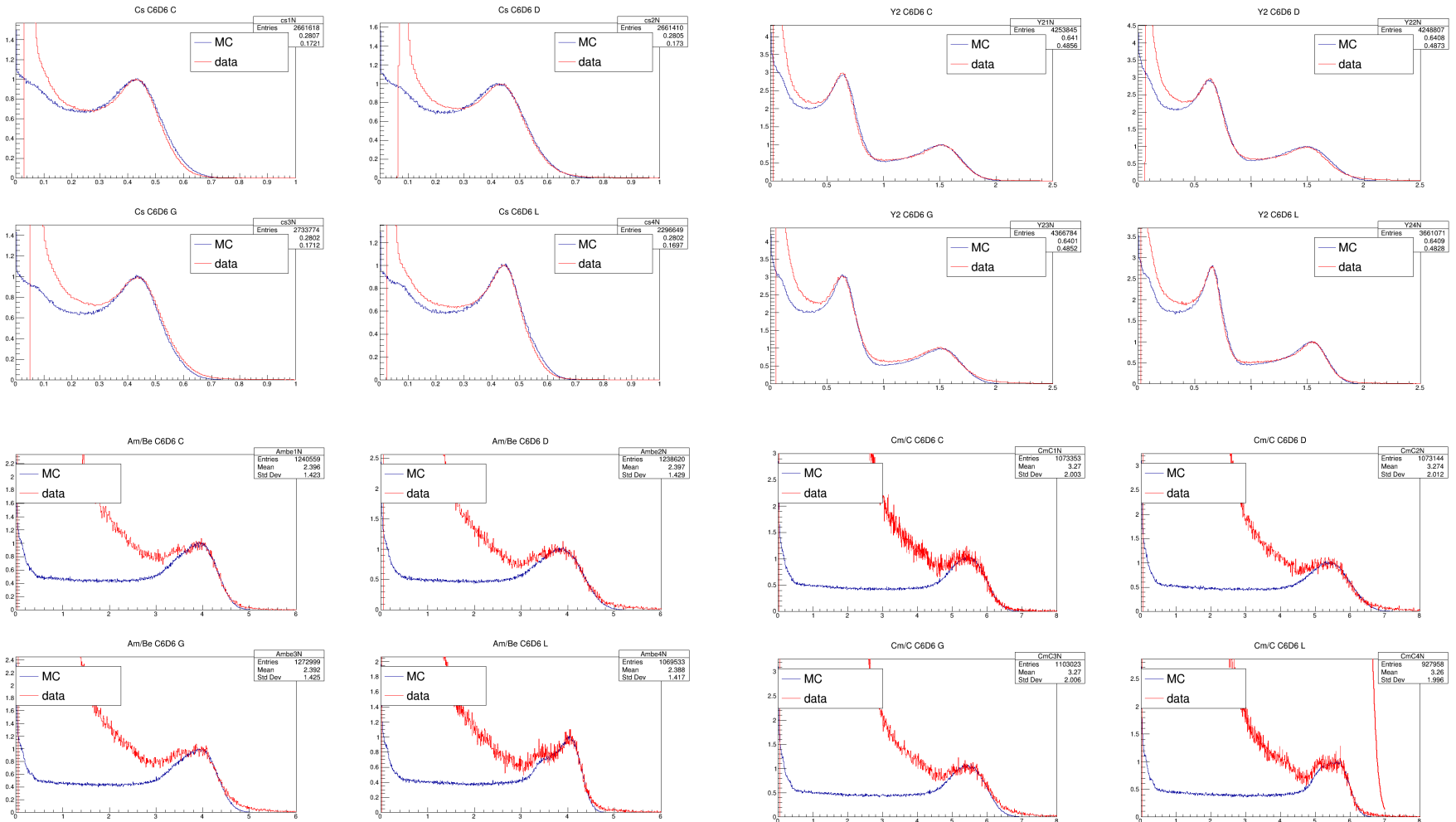
$R(j;E)$ is the simulated detector response to a γ ray of energy E_γ convoluted with the **Experimental Resolution Functions**





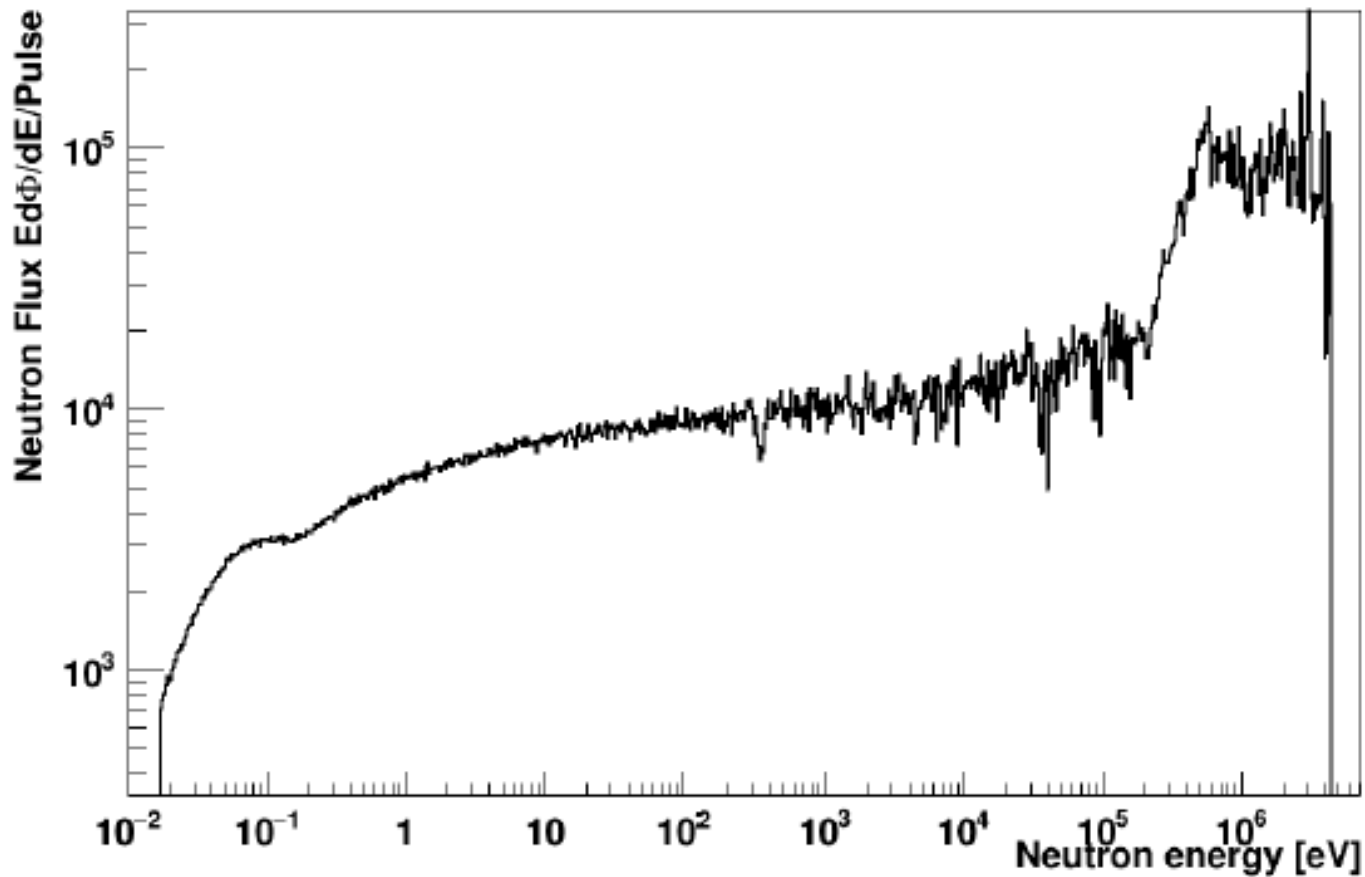
Pulse height weighting technique

Comparison between $R(j;E)$ and the Experimental Detector Response





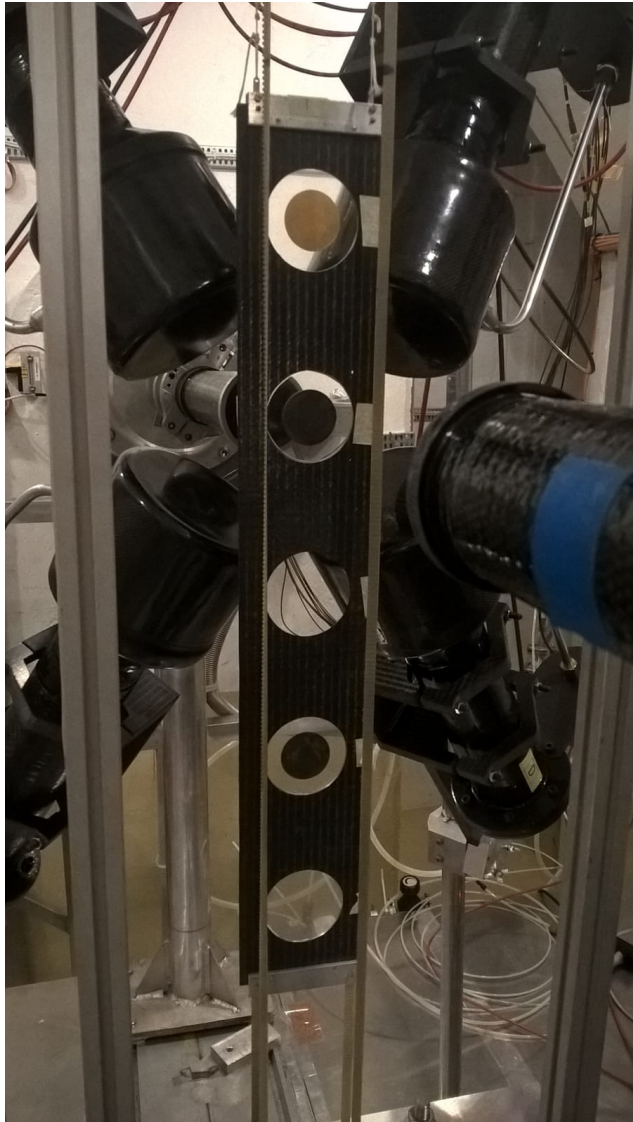
Extracted Flux in EAR1



Extracted flux that will be used for the determination of the capture yield



Background



Gold

Used for the Yield normalization

Lead

Component, caused by scattered in-beam γ rays

Empty

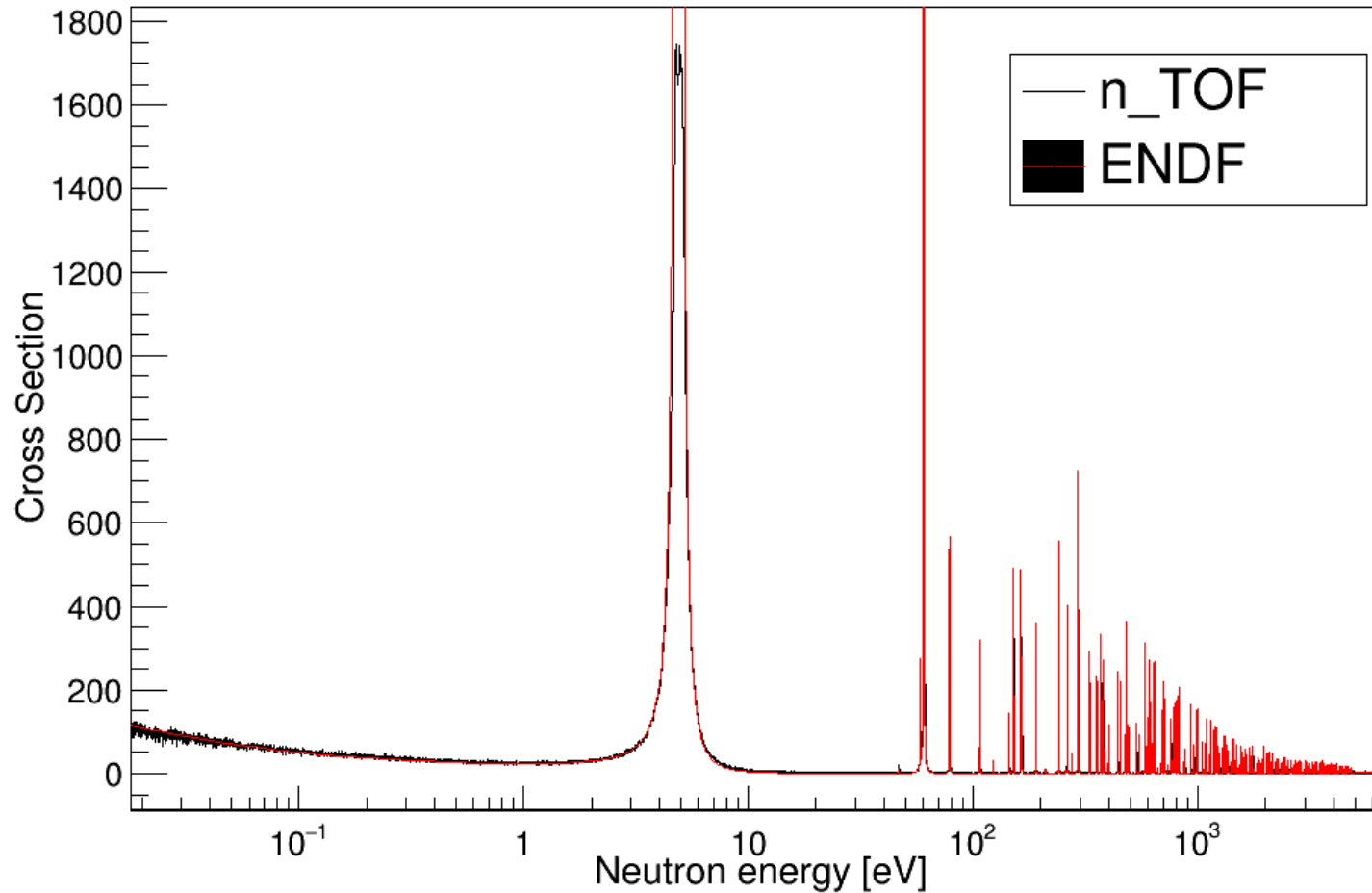
Background related to the neutron beam

Carbon

Neutron background



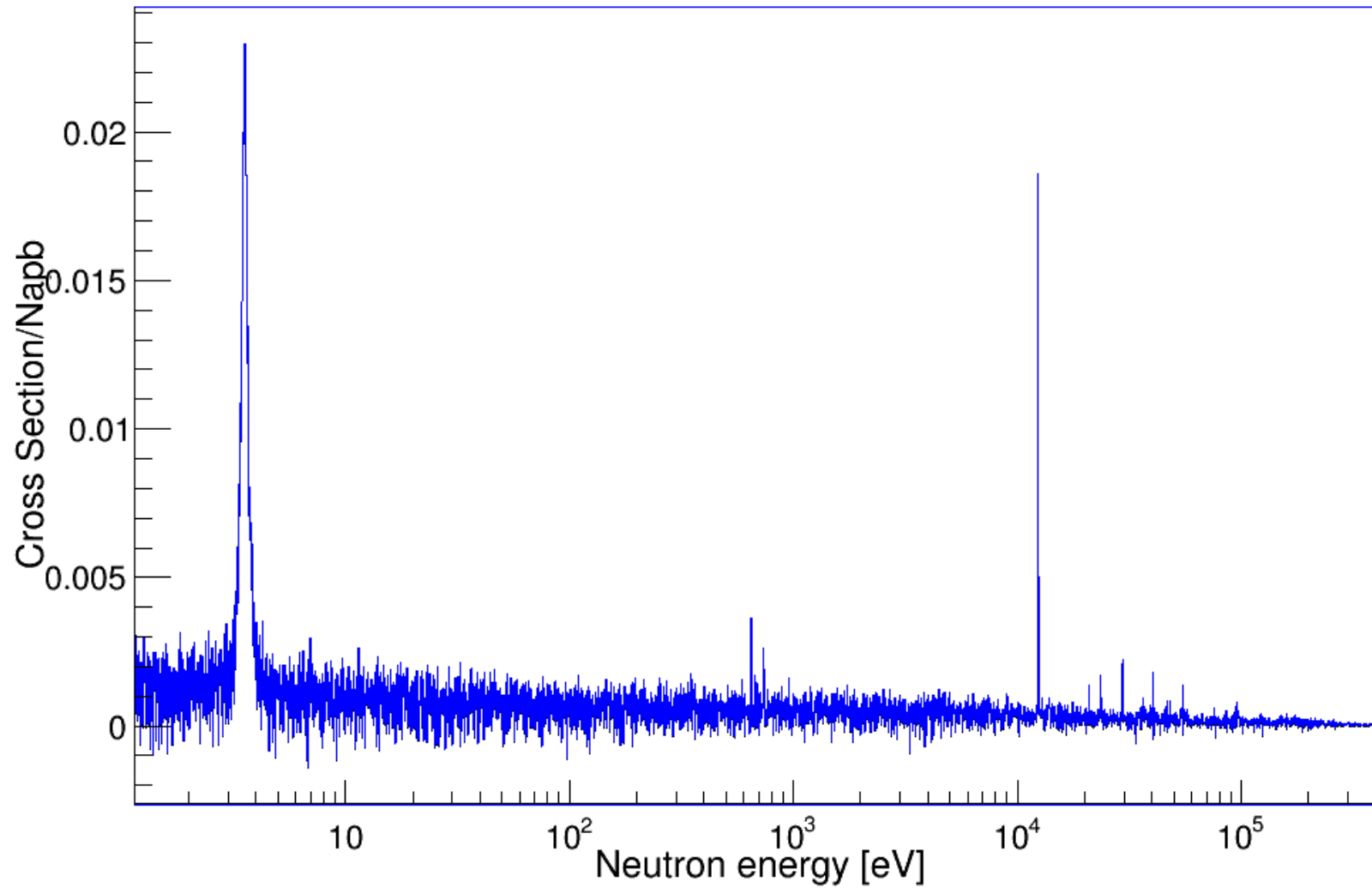
Gold Cross Section



$$\sigma = \frac{\text{Weighted Counts} - \text{Weighted Background}}{\Phi \times N_{apb} \times N \times (S_n + E_{cn} * E_n)}$$

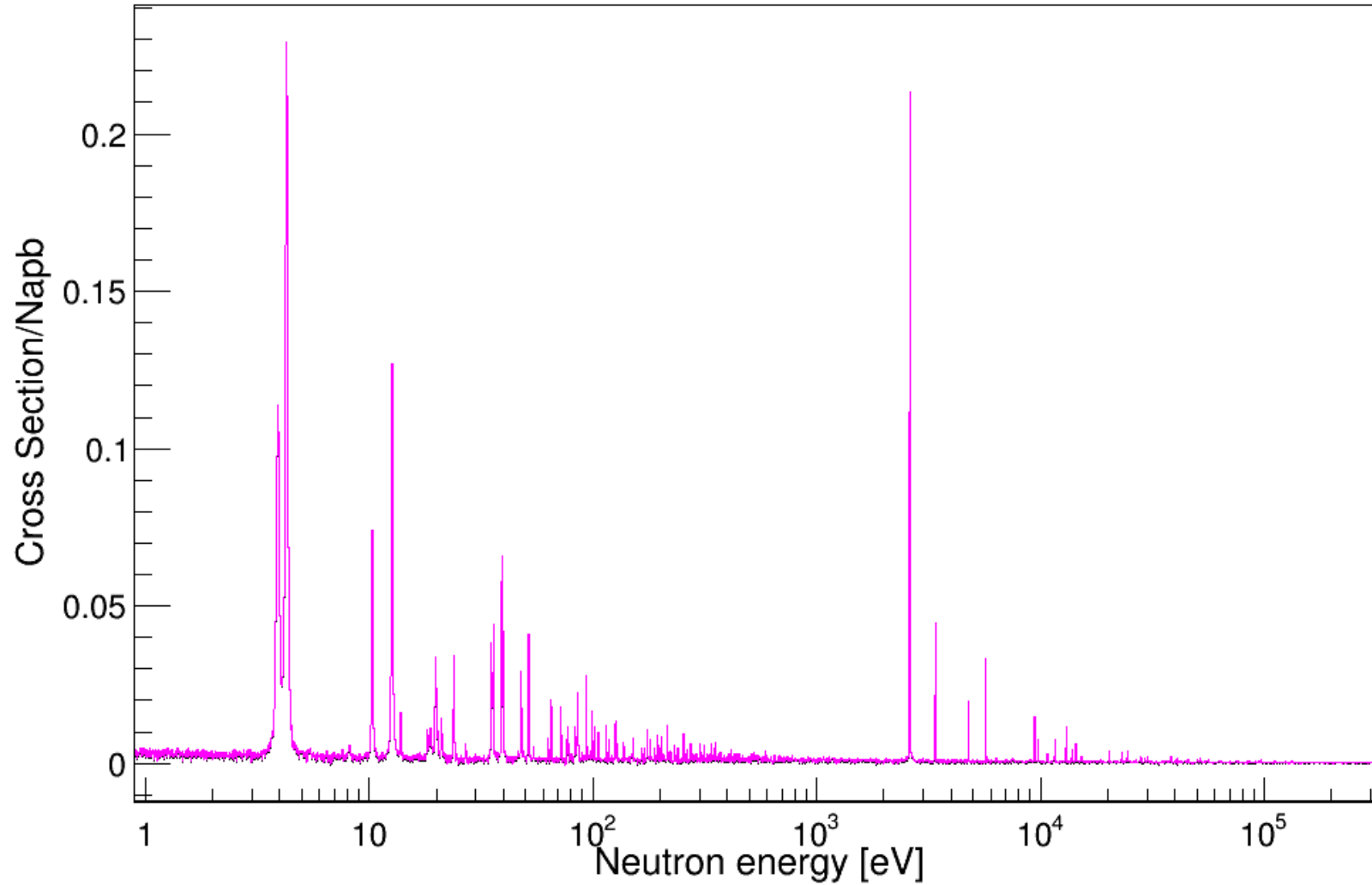


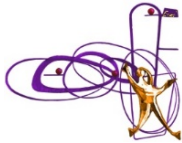
Preliminary results for ^{88}Sr





Preliminary results for ^{89}Y





Publications, Schools and Conferences

Publications

- M. Barbagallo et al., The ${}^7\text{Be}(n,\alpha){}^4\text{He}$ reaction and the Cosmological Lithium Problem: measurement of the cross section in a wide energy range at n TOF (CERN). *Physical Review Letters* 117, 152701 (2016).
- M. Barbagallo et al. ${}^7\text{Be}(n,\alpha)$ and ${}^7\text{Be}(n,p)$ cross section measurement for the Cosmological Lithium problem. *In EPJ Web of Conferences* (2017)
- M. Sabatè Guilarte et al. High-accuracy determination of the neutron flux in the new experimental area n_TOF-EAR2 at CERN" has been accepted for publication in the European Physical Journal A

Schools

- The 9th European Summer School on Experimental Nuclear Astrophysics 17-24 September 2017 Santa Tecla (Italy). (Measurement of the ${}^7\text{Be}(n,p){}^7\text{Li}$ cross section in EAR2@n_TOF for the Cosmological Lithium Problem)

Conferences

- IX Incontro dei Gruppi Italiani di Astrofisica Nucleare Teorica e Sperimentale Sezione INFN di Bologna Dipartimento di Fisica e Astronomia INAF 5-6 October 2017.
- (Measurement of the ${}^7\text{Be}(n,p){}^7\text{Li}$ cross section in EAR2@n_TOF for the Cosmological Lithium Problem)



Goals for the second year of Ph.D

- I will complete the analysis of the radiative neutron capture cross section of ^{89}Y and ^{88}Sr ;
- I will write the thesis;
- I'm planning to spend few months more at CERN to participate in all other measurements that will be performed at n_TOF, helping with the experimental apparatus, the data taking and analysis;
- In 2018 I plan to attend a few conferences where to present my results, and a school in Nuclear Astrophysics;



**THANK YOU FOR YOUR
KIND ATTENTION!**